



Mission Design for Deep Space Nano/Micro Spacecraft Utilizing Lunar Orbital Platform-Gateway Opportunities

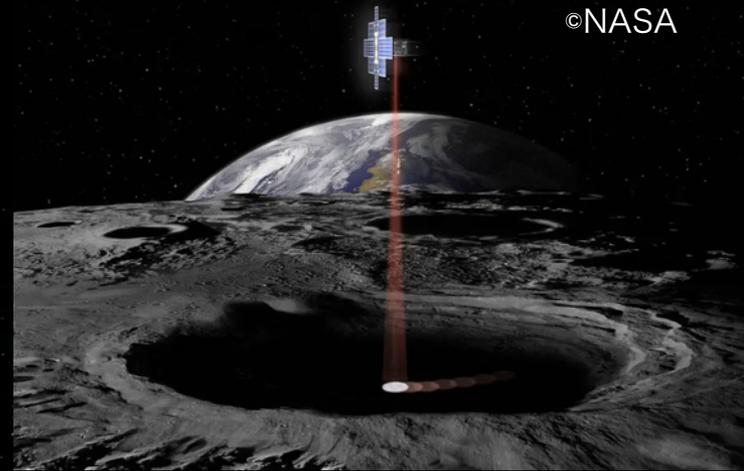
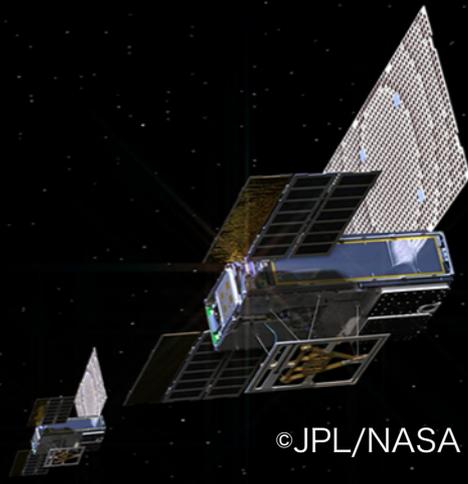
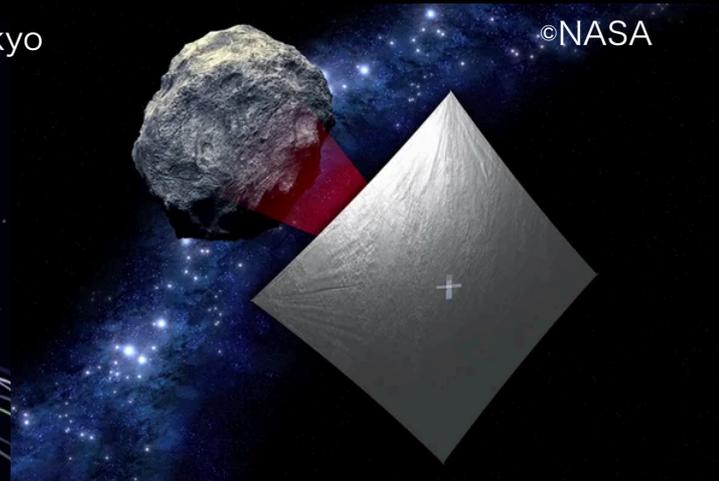
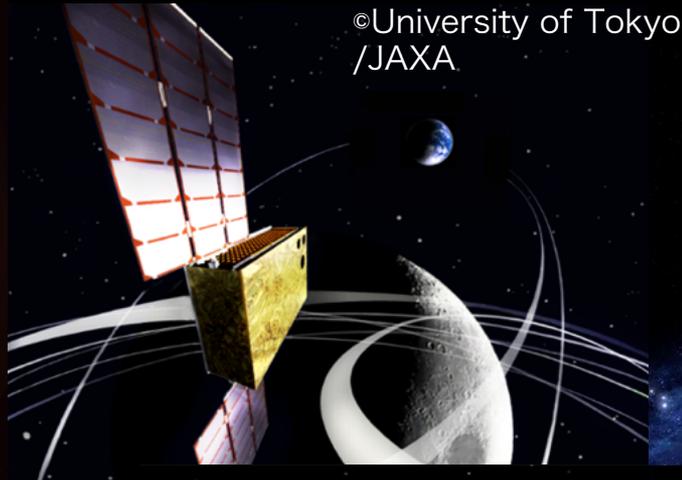
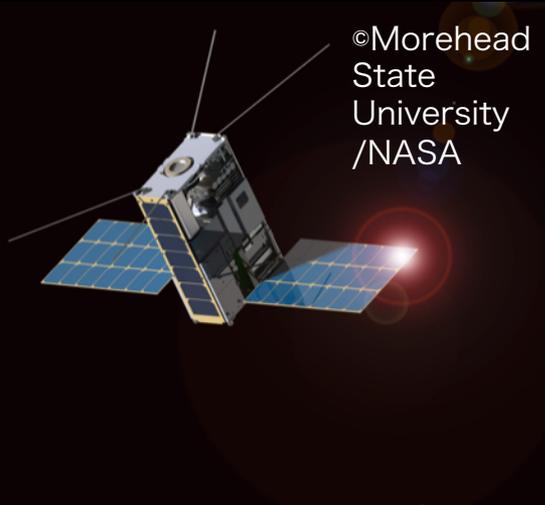
Naoya Ozaki

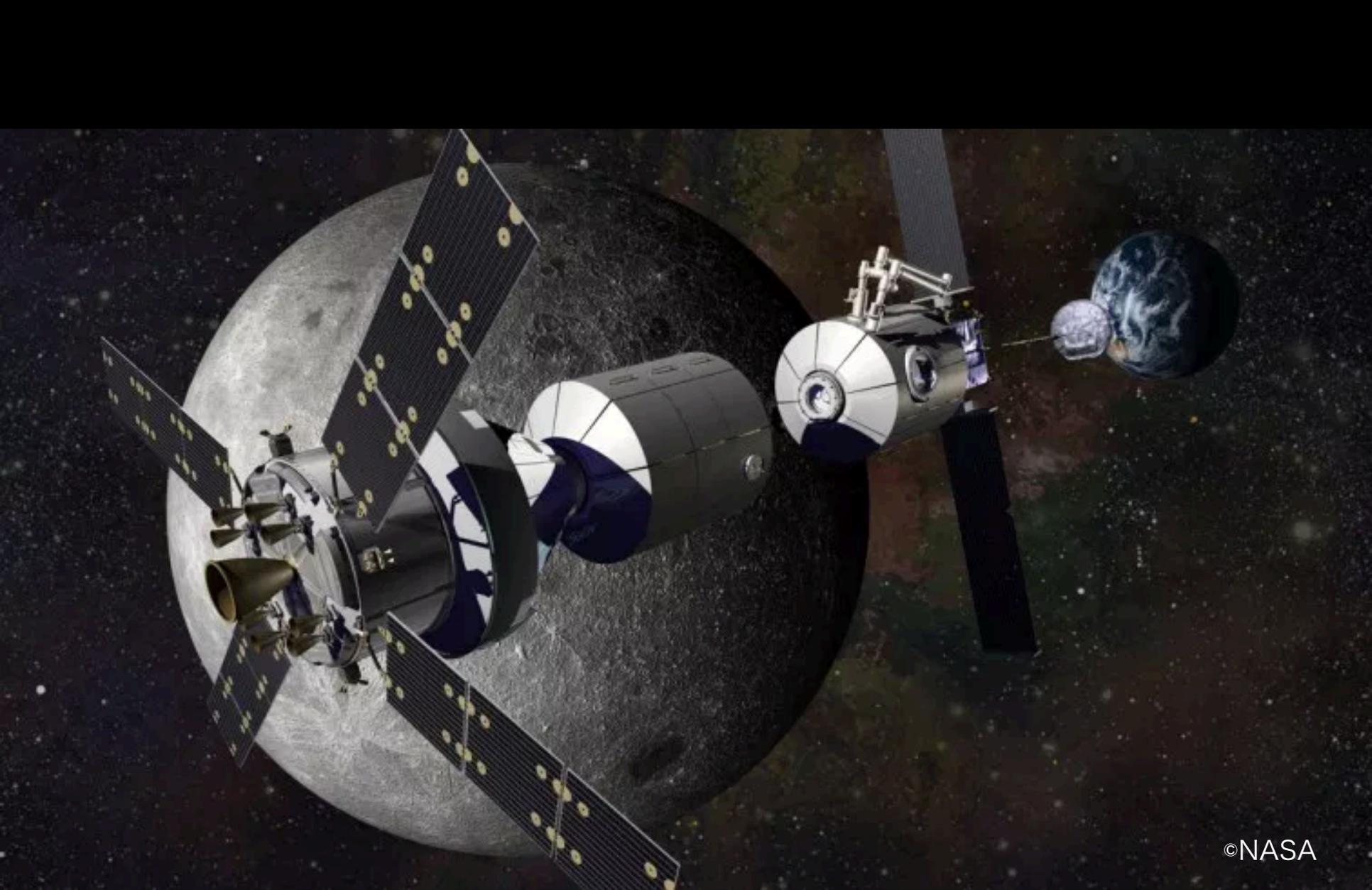
Institute of Space and Astronautical Science,
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NOYCORP

Deep Space CubeSats





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Lunar Orbital Platform-Gateway (In 2020s)

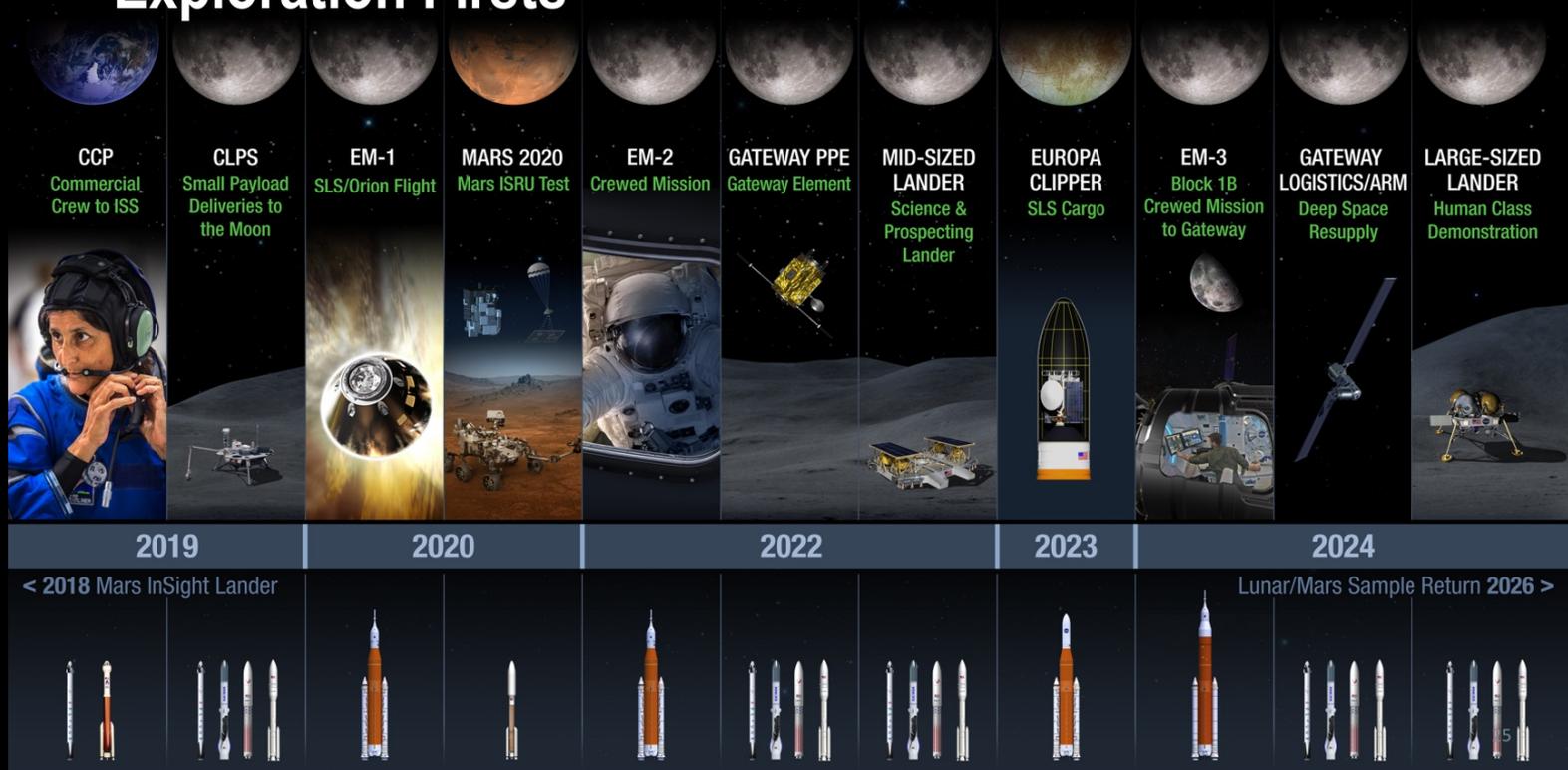
LOP-G Related Launch Opportunities

Starting from NASA's Artemis-1, we can expect more than 10 CubeSats are launched to deep space every year.

(Launch for LOP-G Construction, Resupply, etc...)

Exploration Firsts

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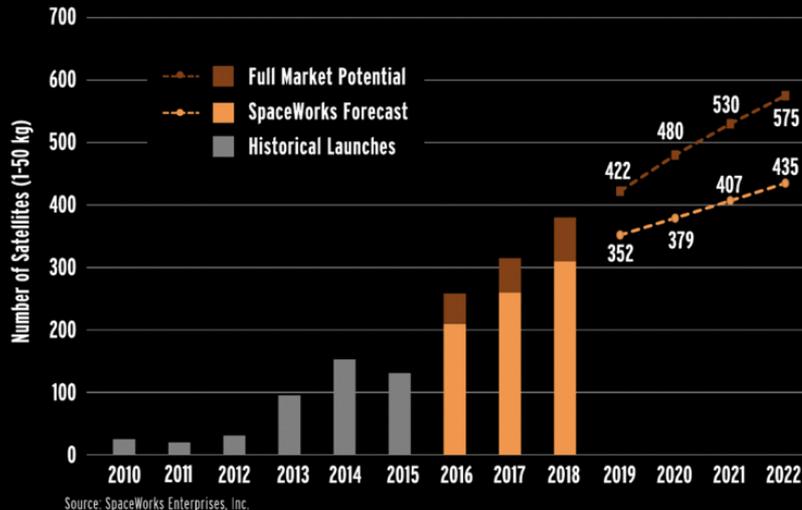


NASA Updates Lunar Gateway Plans, NASA Spaceflight.com (Accessed on March 16, 2019)
<https://www.nasaspaceflight.com/2018/09/nasa-lunar-gateway-plans/>

A New World Opened by LOP-G and Nano/Micro Spacecraft

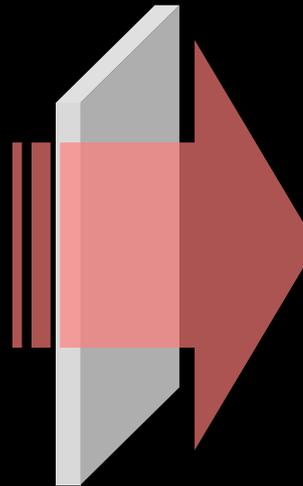
Innovation in Earth satellite (Low cost, short lead time)

- Explosion in numbers
- Frequent missions
- Expansion of stakeholders (Startups, universities, etc)

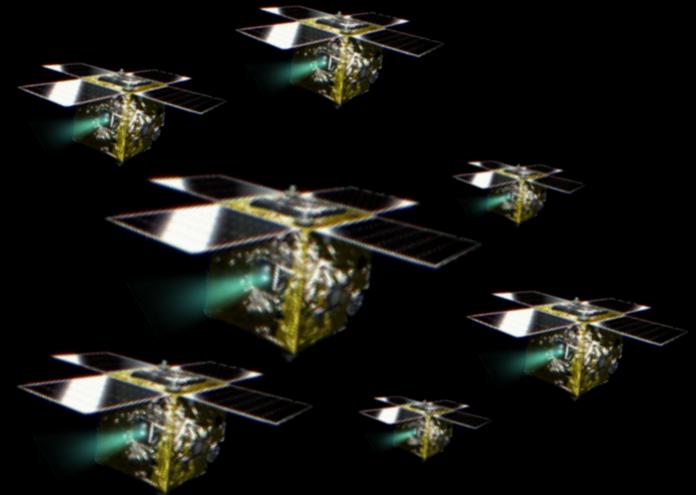


Number of Satellite Launched For A Year

LOP-G

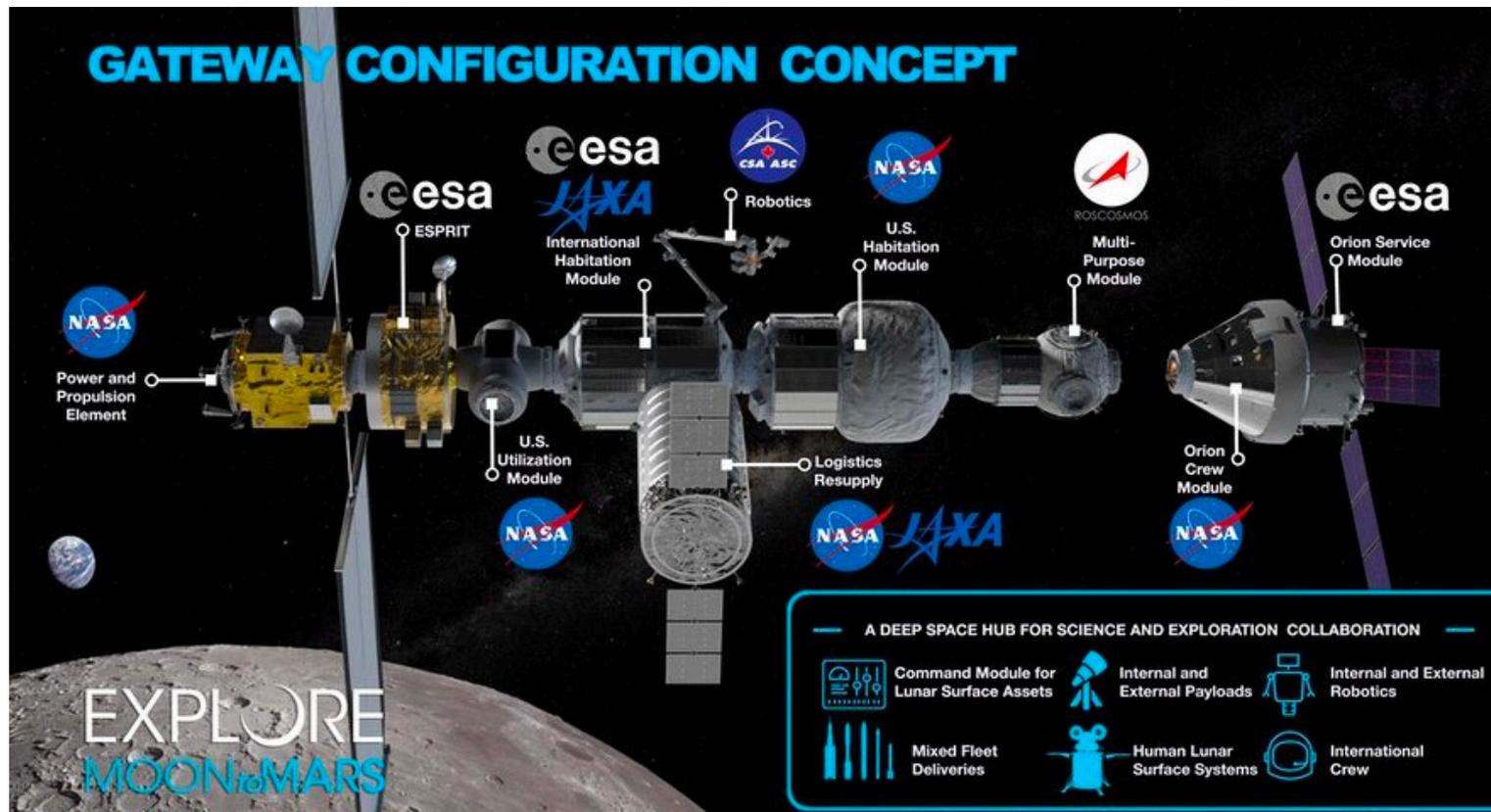


Similar innovation
will happen in
deep space
missions!!



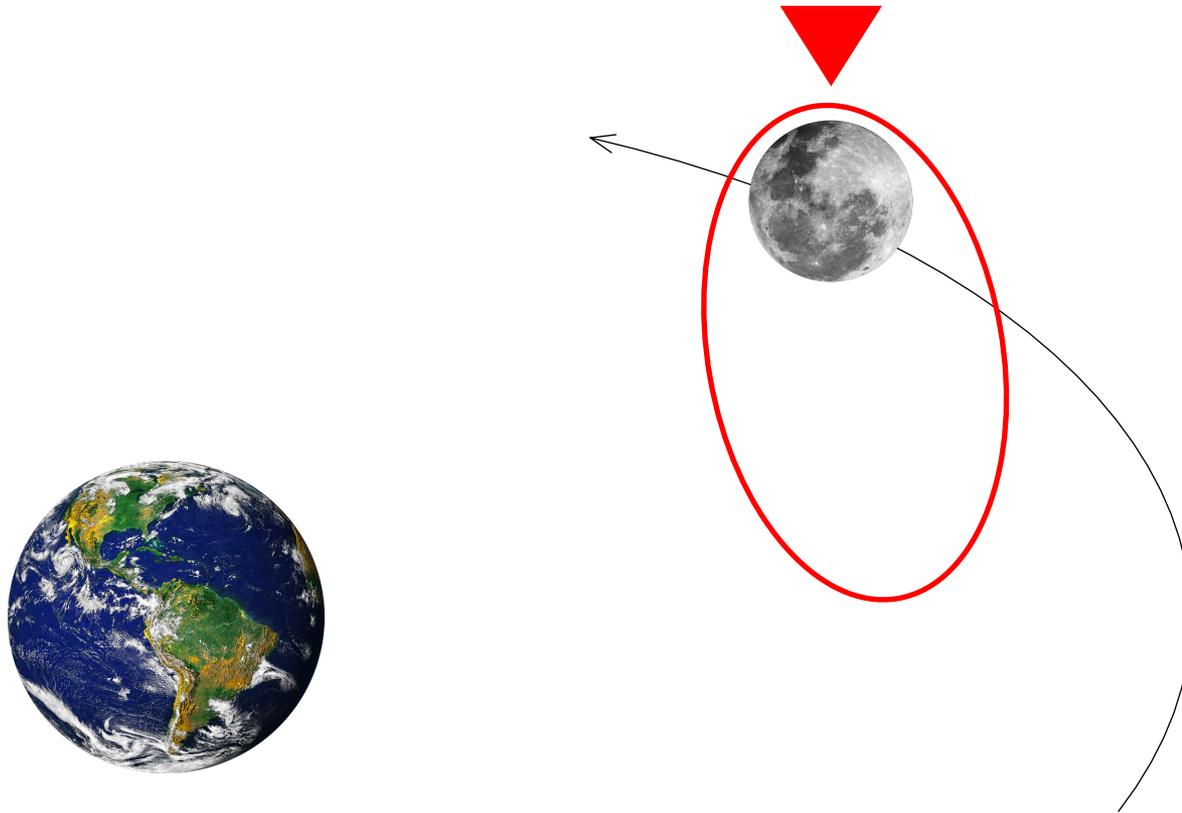
What is Lunar Orbital Platform-Gateway (LOP-G)??

Lunar Orbital Platform-Gateway (LOP-G) is a planned **space station in lunar orbit**. **NASA's Artemis program** plays a major role to develop the Gateway in collaboration with commercial and international partners: ESA, JAXA, CSA, Roscosmos, etc.



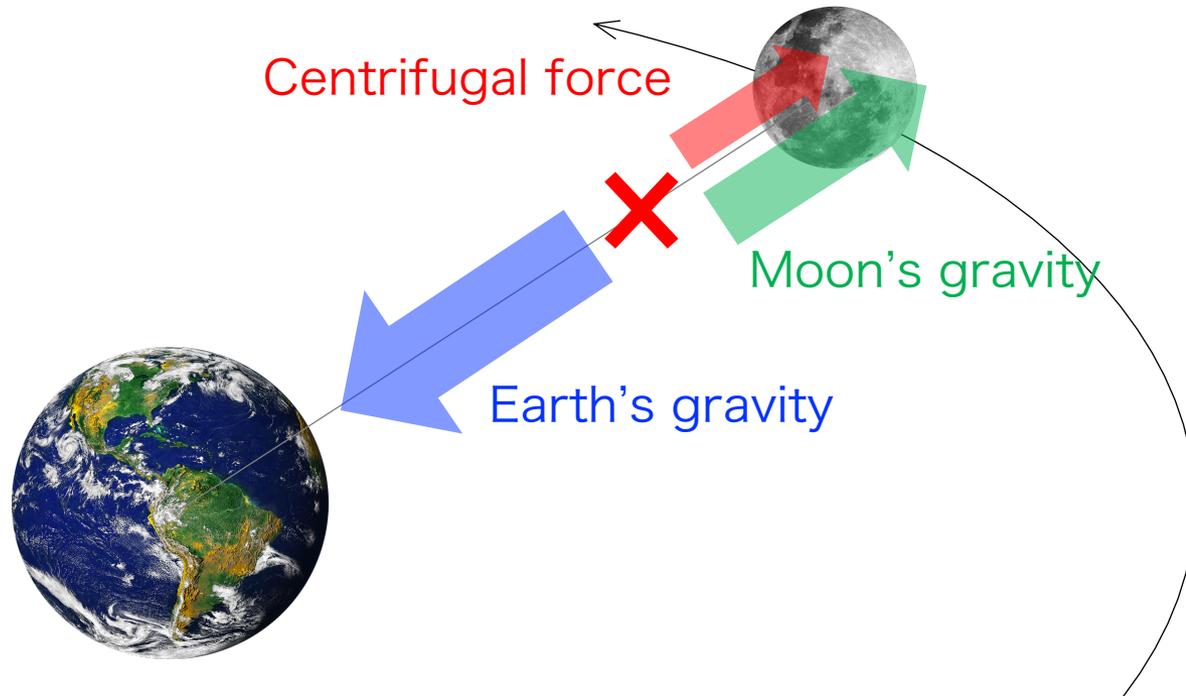
Where is the Gateway??

Near Rectilinear Halo Orbit



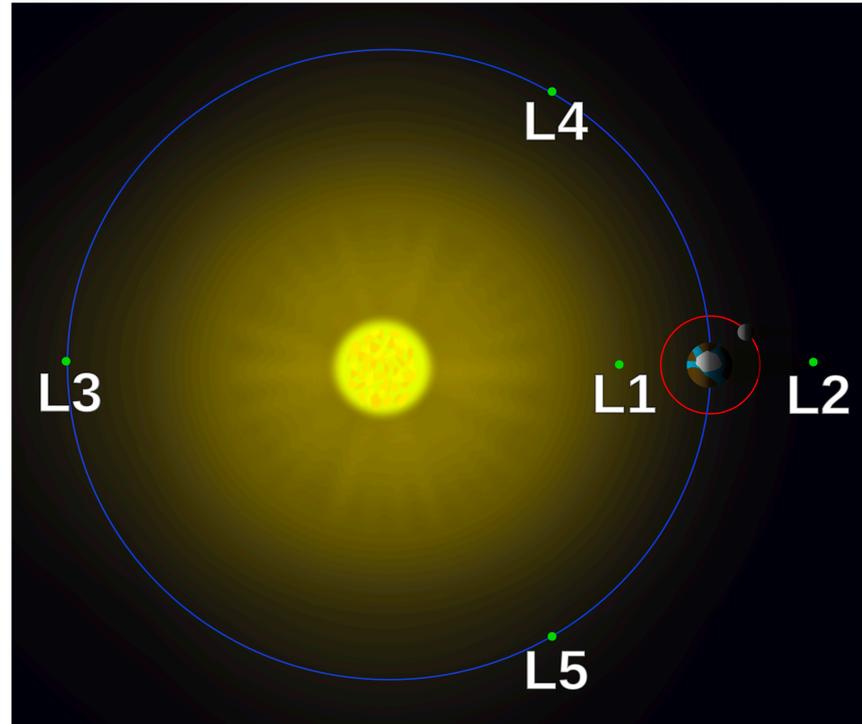
A type of Halo orbits under the Earth and Moon gravity.

Where is the Gateway??



The equilibrium point where the earth's gravity, the moon's gravity, and the centrifugal force balance each other is called **the Lagrange point**.

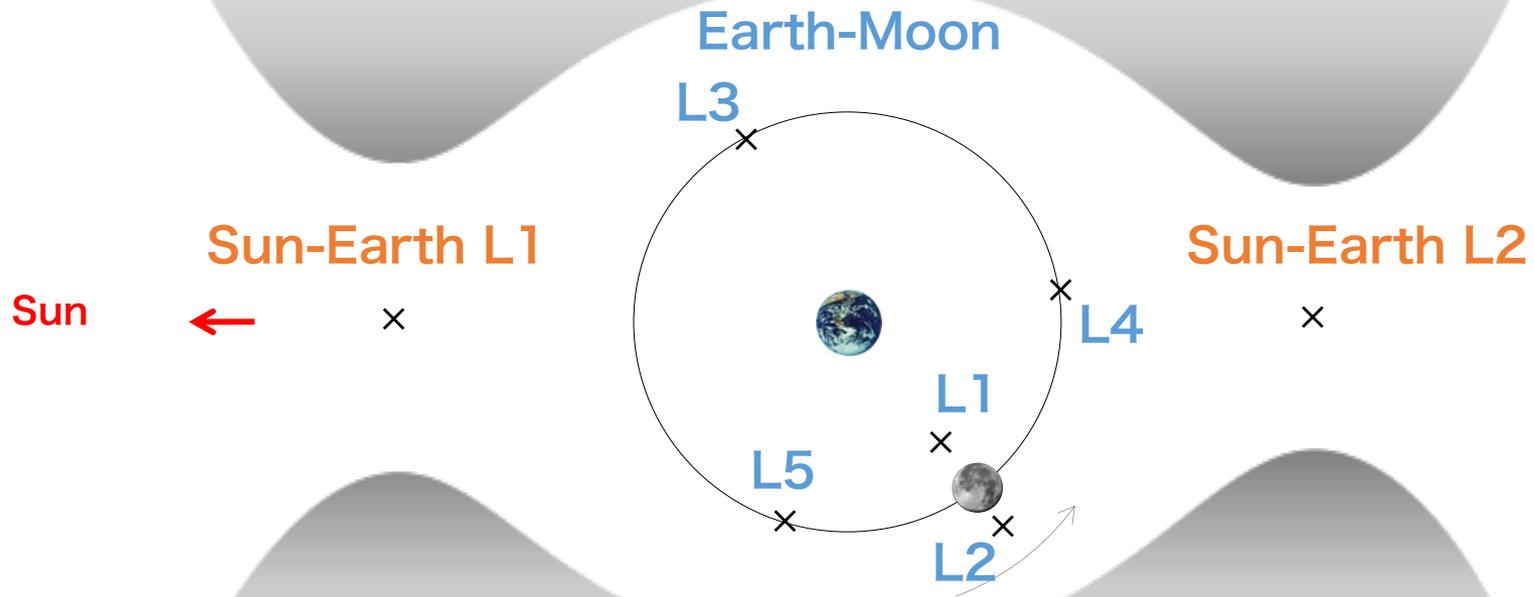
Lagrange Points



Five types of Lagrange points exist in each three-body dynamical system.

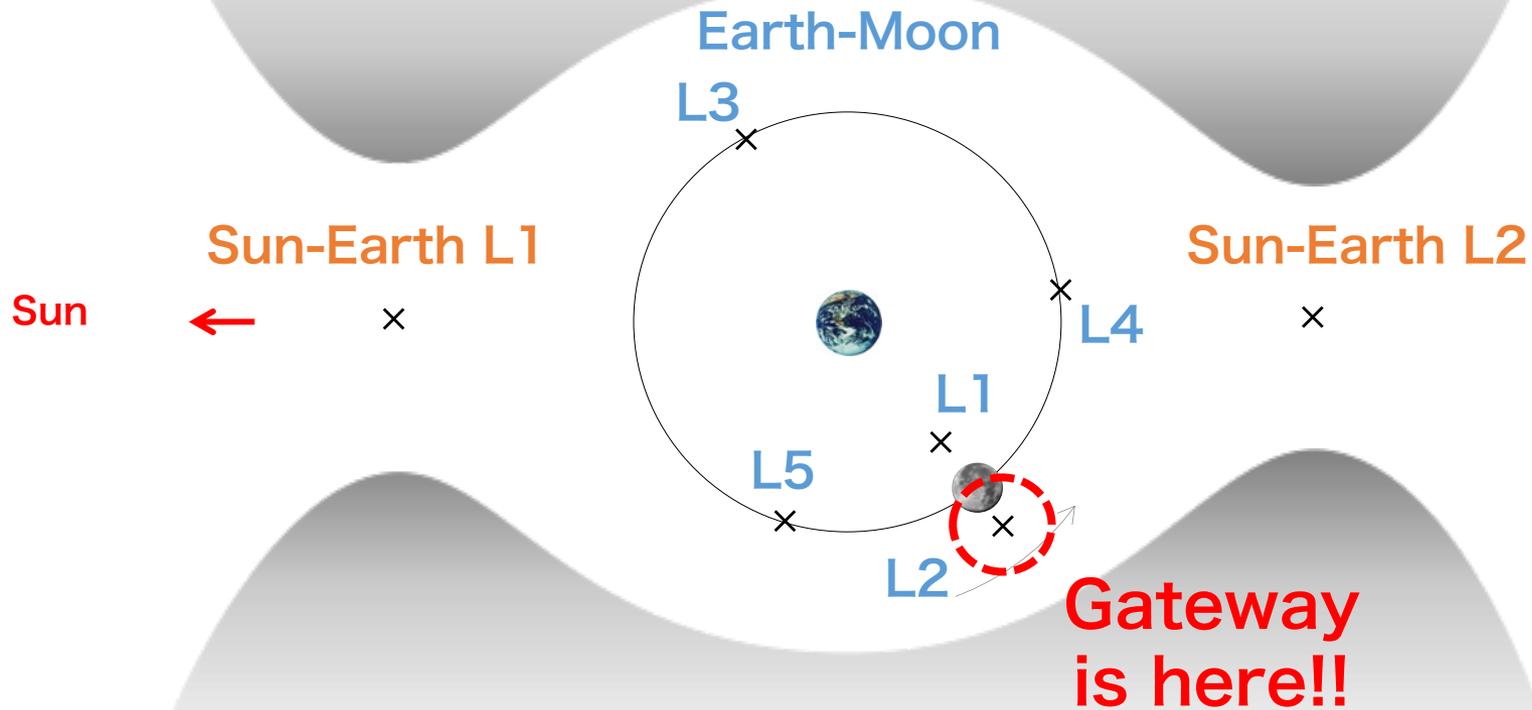
Ex) Earth-Moon L2 Lagrange point
Sun-Earth L1 Lagrange point

Geometry of Lagrange Points



Sun-Earth line fixed rotational frame.

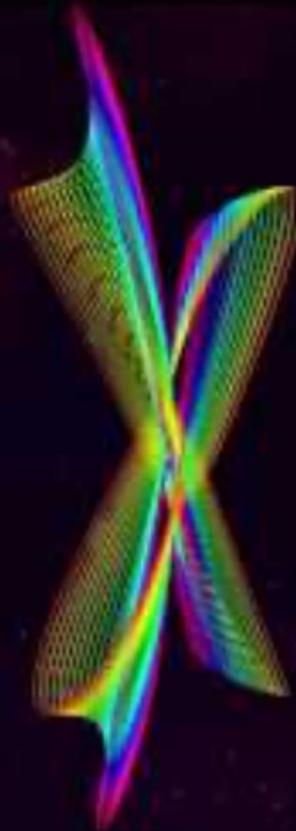
Geometry of Lagrange Points



Sun-Earth line fixed rotational frame.

Near Rectilinear Halo Orbit

P1: Earth
P2: Moon
Radius P2: 0.0045 Scaling: 1.0
JC: 3.0594349784036745



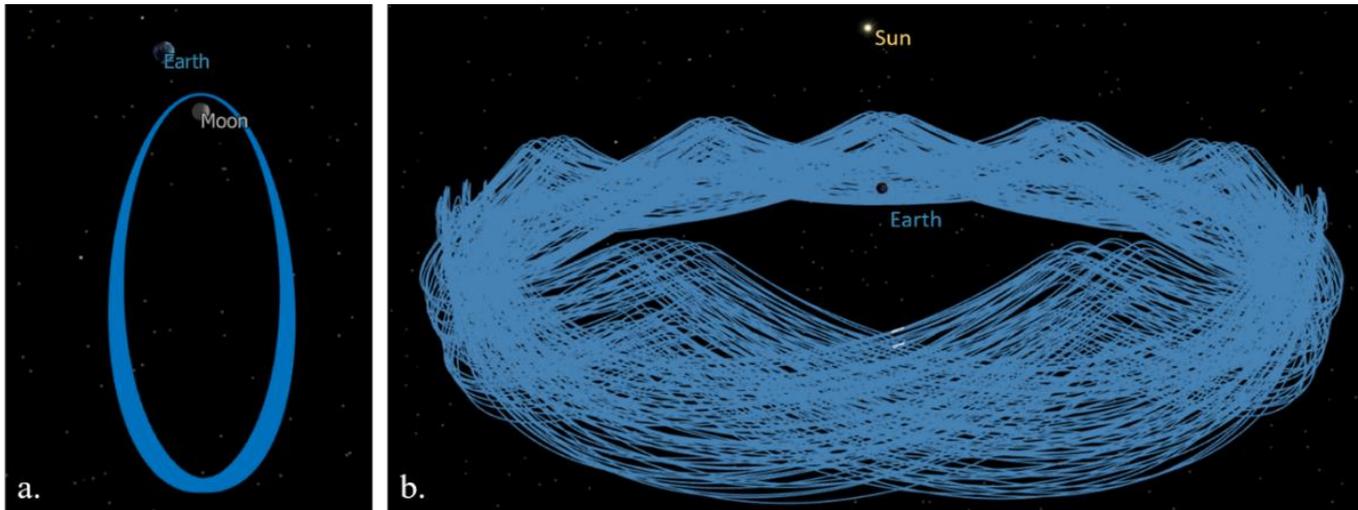
Jacobi Constant: 2.99957
Y Amplitude: 3.735951E04 km
Z Amplitude: 7.518863E04 km
Period: 0008.9130 days
Close Approach: 7.071249E05 km

The Northern and Southern L₁ and L₂ NRHOs are periodic in the Circular Restricted 3-Body Model, and can be transitioned into quasi-periodic orbits in a higher fidelity model

Near Rectilinear Halo Orbit (NRHO)

Characteristics of NRHO

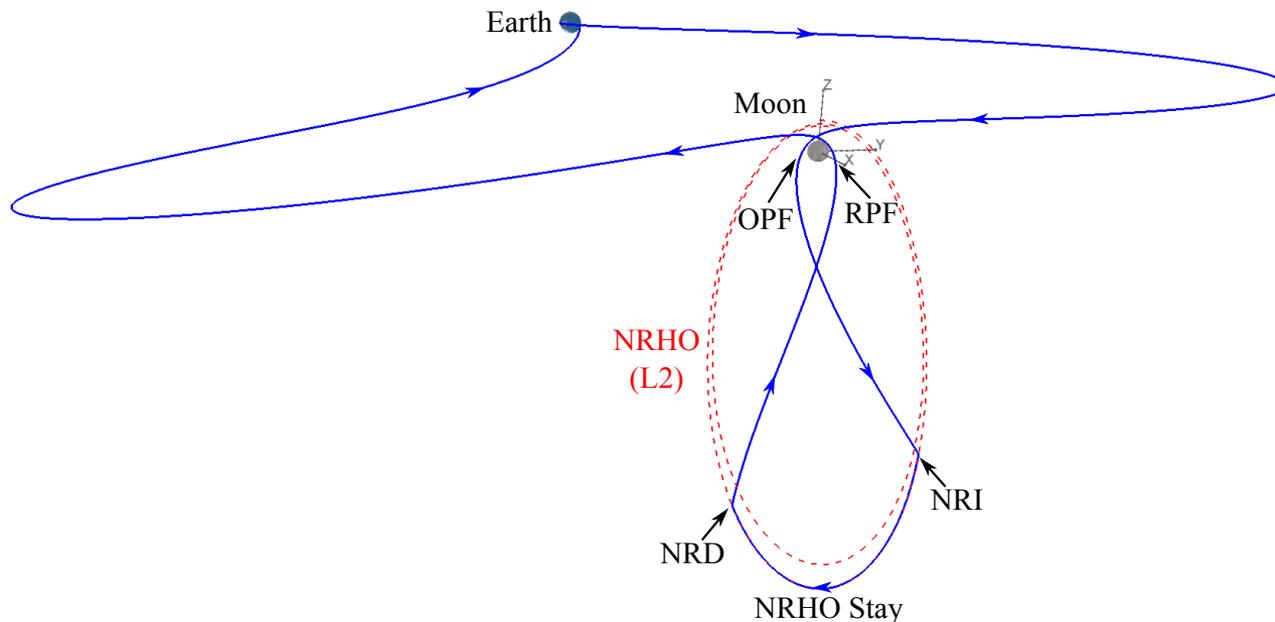
- **Geometric relationship to the Earth is always the same**
- Halo orbits are unstable (**easily reachable/escapable**) and spacecraft cannot stay in the orbits without station keeping maneuvers, but NRHO is less stable than general Halo orbits.
- Gateway will be constructed in 9:2 synodic resonant NRHO (SRHO), where **the station never experiences eclipse**.
- For 9:2 SRHO, the perilune altitude is 1458-1820km, and the apolune altitude is 68267-70112km.



Launch Condition for Gateway Construction Opportunity

For gateway construction opportunities, the spacecraft is expected to be launched into **lunar transfer trajectory** (as was the case with Artemis 1).

Using lunar swing-by on this orbit, the spacecraft can fly to **interplanetary space (to asteroid, Mars, ...)**, **periodic orbits in Lagrange points**, and so on.



Mission Utilizing Gateway Construction Opportunity

Launch & Separation



Lunar transfer trajectory

~1 week

Trajectory correction ΔV
~20m/s



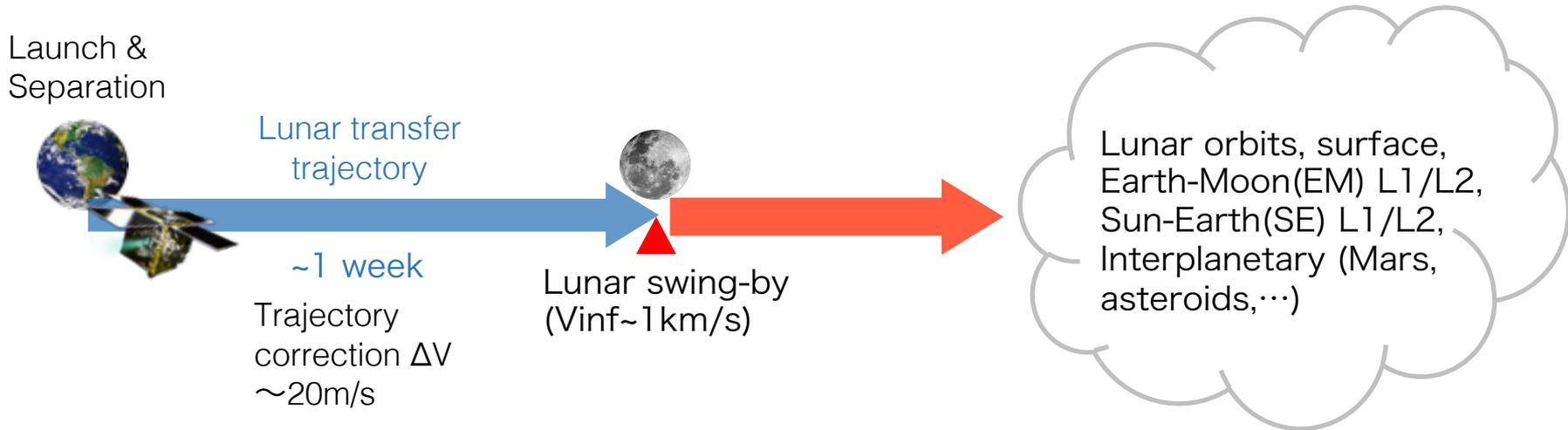
Lunar swing-by
(V_{inf} ~1km/s)



Lunar orbits, surface,
Earth-Moon(EM) L1/L2,
Sun-Earth(SE) L1/L2,
Interplanetary (Mars,
asteroids,···)

Possible mission scenario for early Artemis opportunity (such as Artemis 2)

Mission Utilizing Gateway Construction Opportunity



1) To Sun-Earth L1/L2: Ballistic transfer without ΔV

2) To Lunar Surface:

3) To Lunar orbits, Earth-Moon L1/L2:

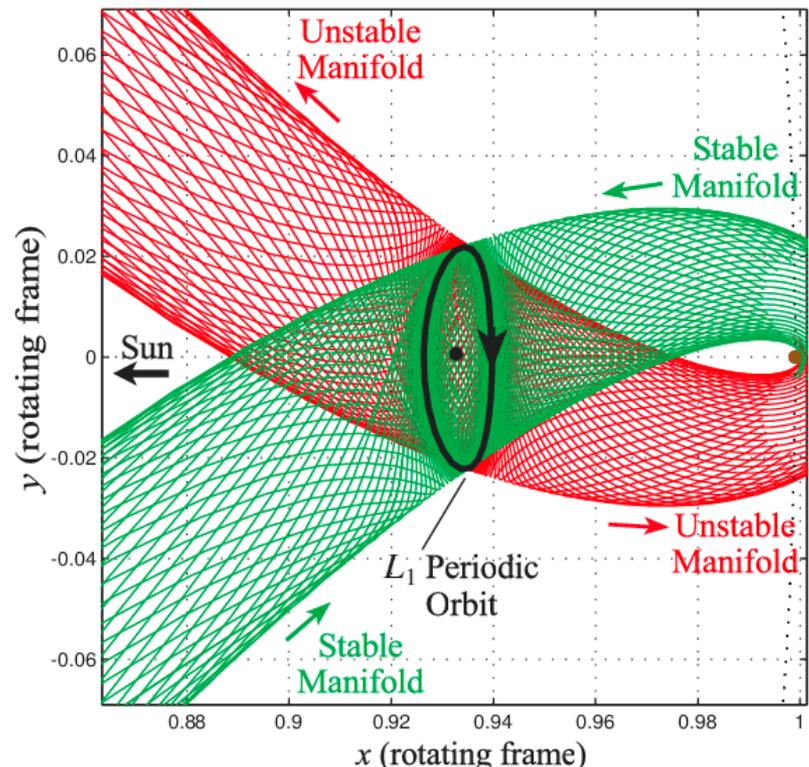
4) To Interplanetary (Mars, asteroids):

1) Transfer to Sun-Earth L1/L2 Points

There is a manifold structure in periodic orbits around Lagrange points.

When a small ΔV disturb at each difference phase on the periodic orbit, the spacecraft leaves the periodic orbits on a group of orbits called **the unstable manifold**.

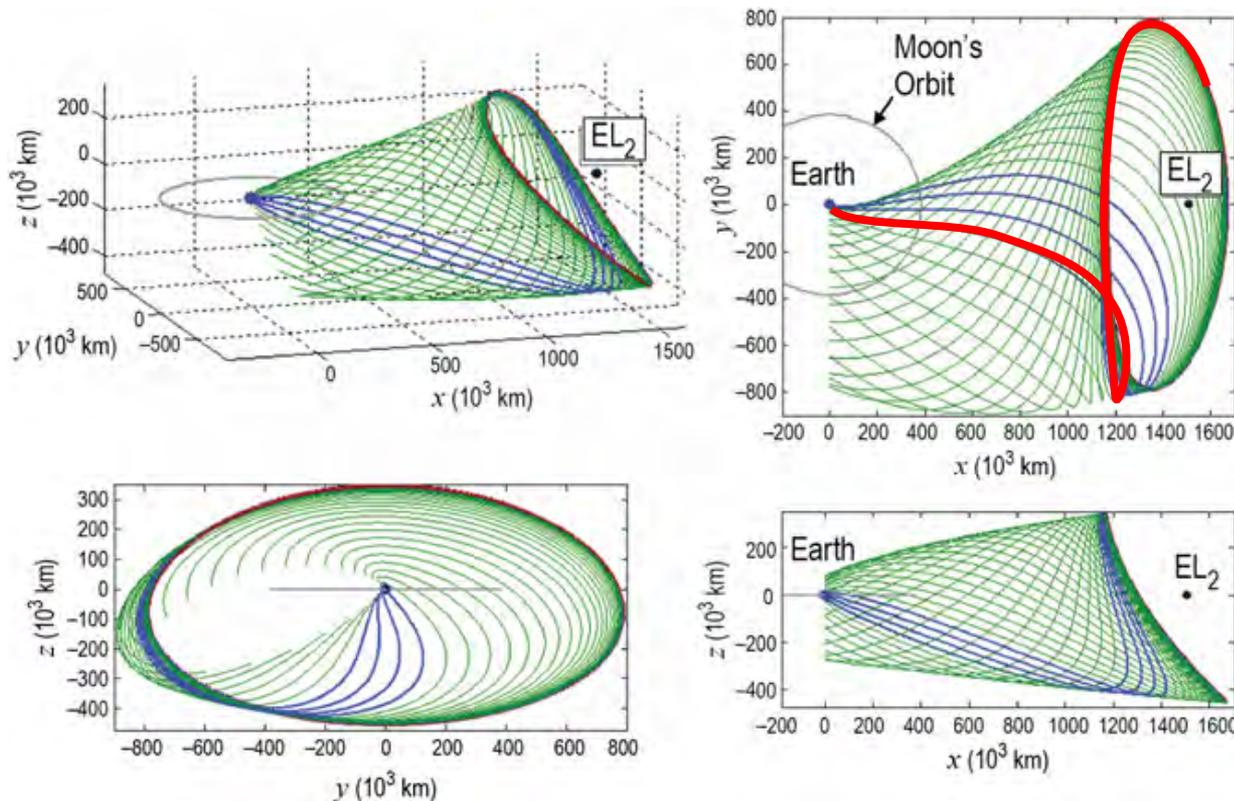
Because of the symmetry, the spacecraft can asymptotically approach a period orbit by riding on a group of orbits called **the stable manifold**.



(W.S., Koon, M.W., Lo, et al. , *AAS*, 2000)

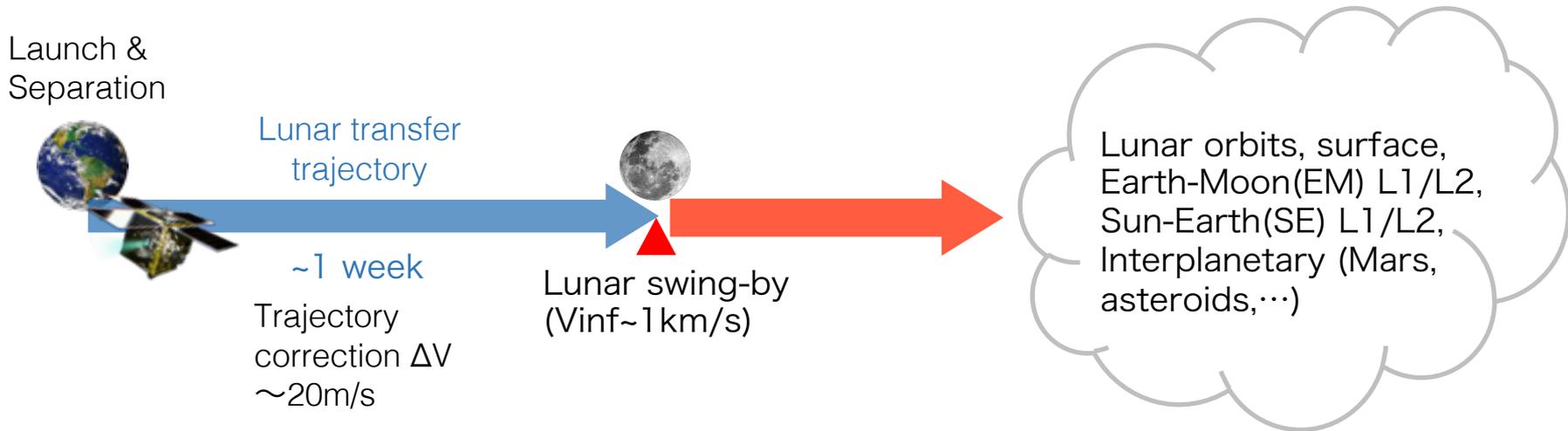
1) Transfer to Sun-Earth L1/L2 Points

By selecting a single trajectory on a stable manifold, the spacecraft can ballistically transfer to the periodic orbit around the Lagrange point.



(J.S.Parker and R.L.Anderson, *JPL*, p.97, 2013)

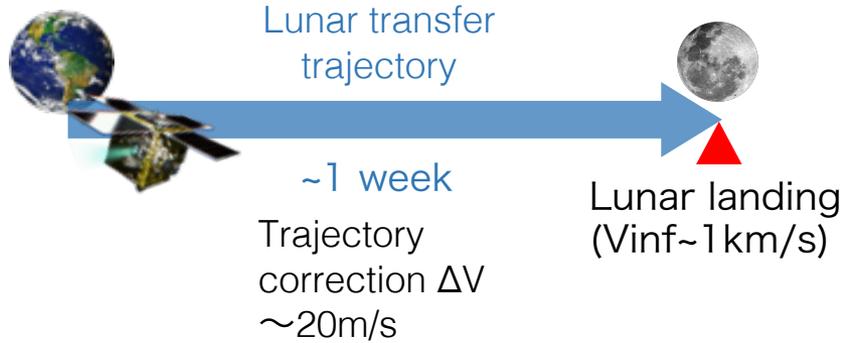
Mission Utilizing Gateway Construction Opportunity



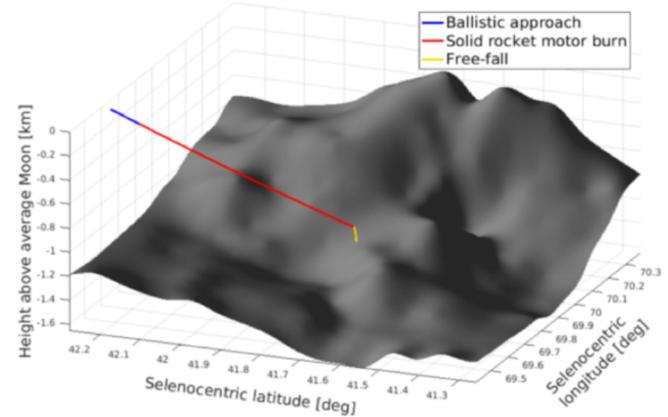
- 1) To Sun-Earth L1/L2: Ballistic transfer without ΔV
- 2) To Lunar Surface: Landing with 2.5km/s ΔV
- 3) To Lunar orbits, Earth-Moon L1/L2:
- 4) To Interplanetary (Mars, asteroids):

2) Landing on Lunar Surface

Launch & Separation



OMOTENASHI Trajectory



(S. Campagnola, et al., IEEE, 2019)

It is possible to estimate the landing ΔV by assuming a two-body problem (patched conics)

In the vicinity of the moon, vis-viva equation (orbital-energy-invariance law) gives

$$\frac{1}{2}v^2 - \frac{GM_M}{r} = \frac{1}{2}v_{\infty}^2$$

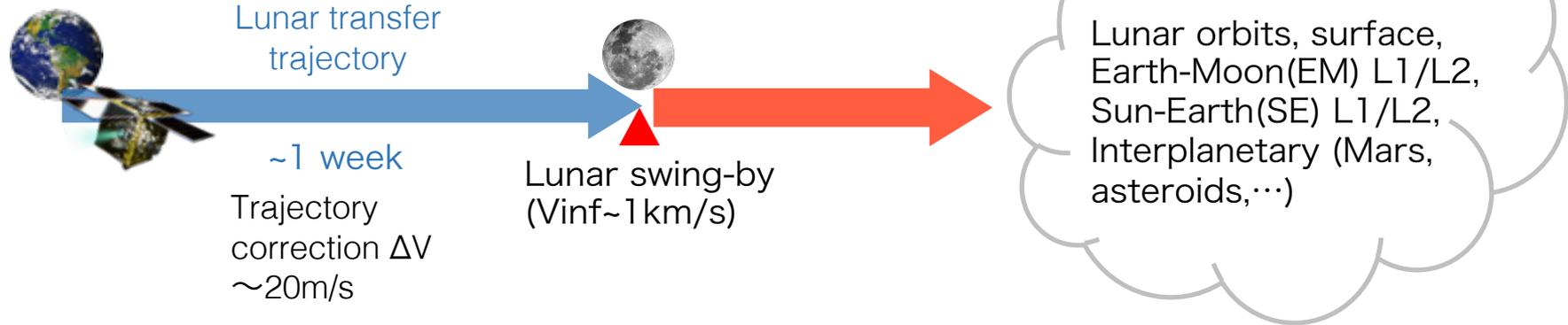
Suppose that $v_{\infty} = 0.82\text{km/s}$ (example of Artemis 1) and $r = r_M$ (lunar radius), the velocity is

$$v = 2.514 \text{ km/s}$$

In order to land on the moon, we need to cancel this velocity, i.e., **$\Delta V \sim 2.5\text{km/s}$** .

Mission Utilizing Gateway Construction Opportunity

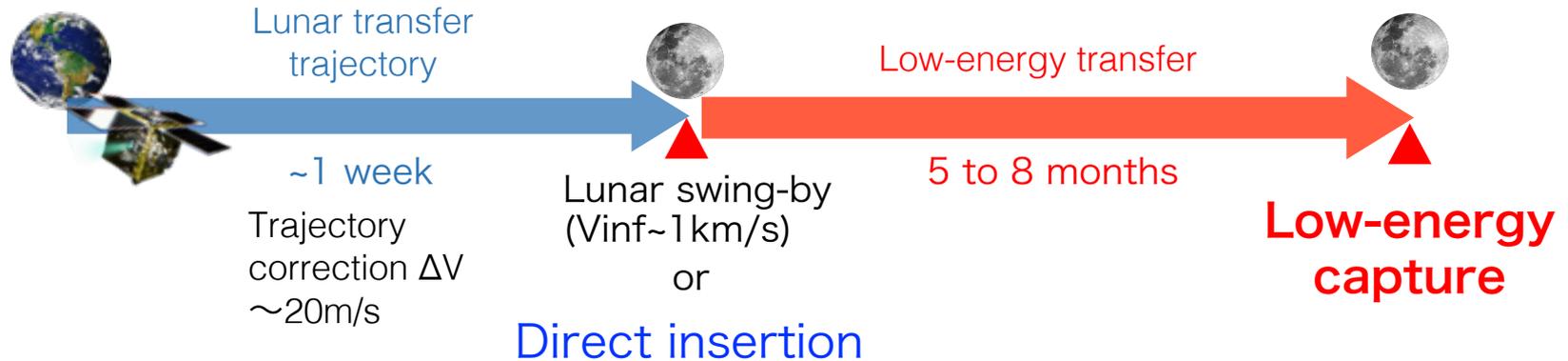
Launch & Separation



- 1) To Sun-Earth L1/L2: Ballistic transfer without ΔV
- 2) To Lunar Surface: Landing with 2.5km/s ΔV
- 3) To Lunar orbits, Earth-Moon L1/L2: Direct insertion or Low-energy transfer/capture by reducing V_{inf} .
- 4) To Interplanetary (Mars, asteroids):

3) Transfer to Lunar Orbits or Earth-Moon L1/L2 points

Launch & Separation

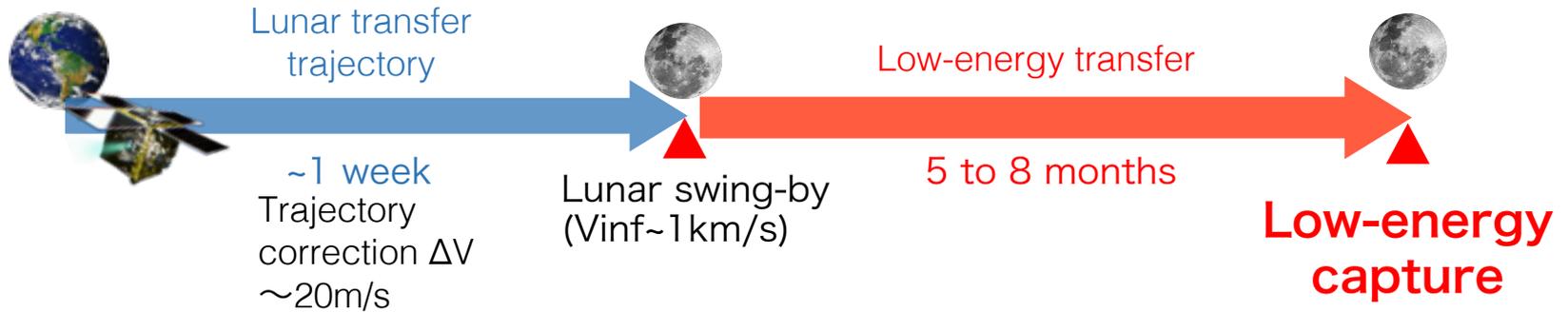


For direct insertion, it is possible to estimate the insertion ΔV by assuming a two-body problem (patched conics). The ΔV is about 0.5-1km/s.

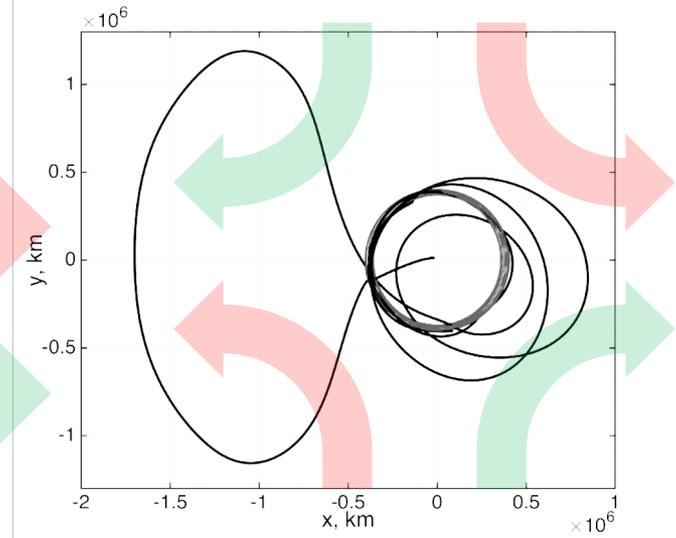
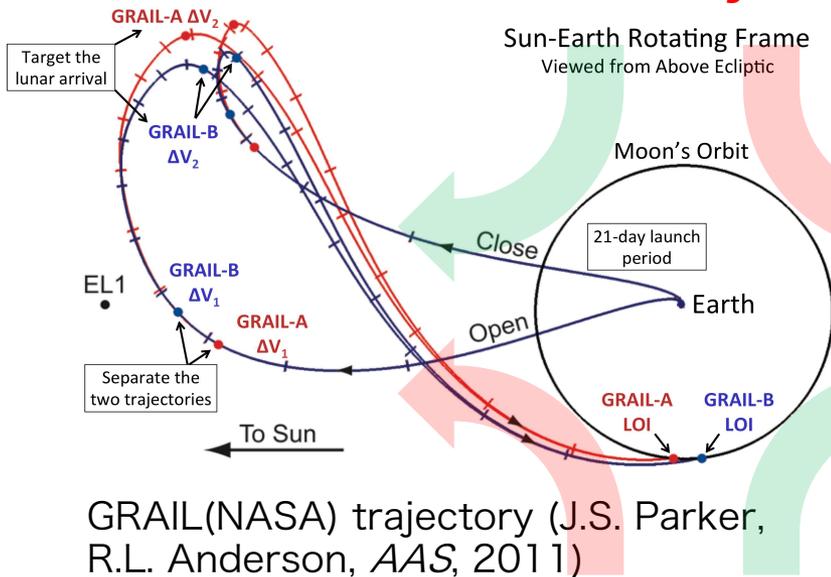
For low-energy transfer/capture, the solar tidal force can effectively reduce V_{∞} . For NRHO or other Halo orbits, only about 10m/s ΔV is required for the insertion.

3) Transfer to Lunar Orbits or Earth-Moon L1/L2 points

Launch & Separation



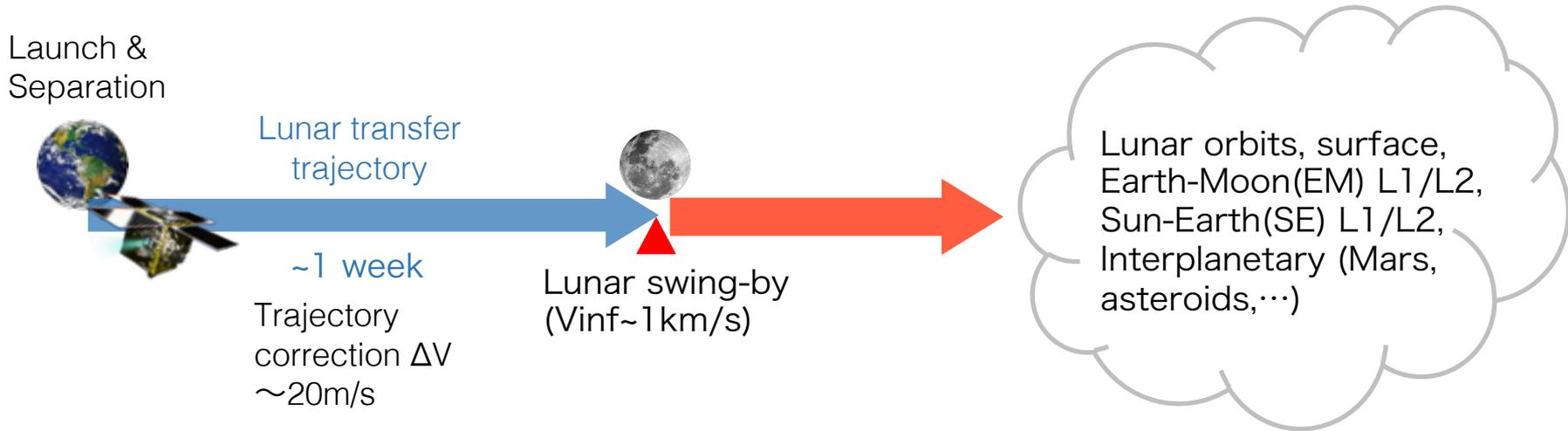
For low-energy transfer/capture, the solar tidal force can effectively reduce V_{inf} without ΔV .



EQUULEUS trajectory (example)

GRAIL(NASA) trajectory (J.S. Parker, R.L. Anderson, AAS, 2011)

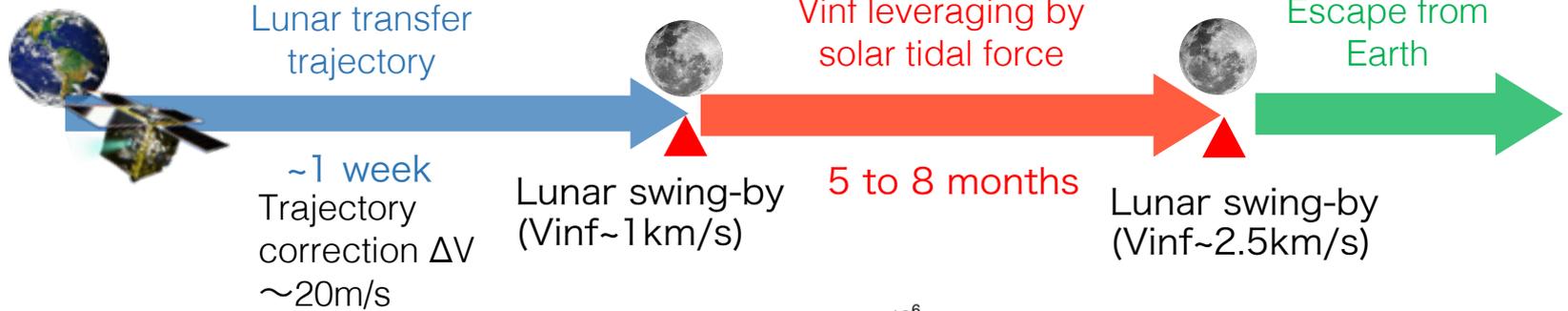
Mission Utilizing Gateway Construction Opportunity



- 1) To Sun-Earth L1/L2: Ballistic transfer without ΔV
- 2) To Lunar Surface: Landing with 2.5km/s ΔV
- 3) To Lunar orbits, Earth-Moon L1/L2: Direct insertion or Low-energy transfer/capture by reducing V_{inf} .
- 4) To Interplanetary (Mars, asteroids): Escaping by leveraging V_{inf} .

4) To Interplanetary Space (Mars, etc)

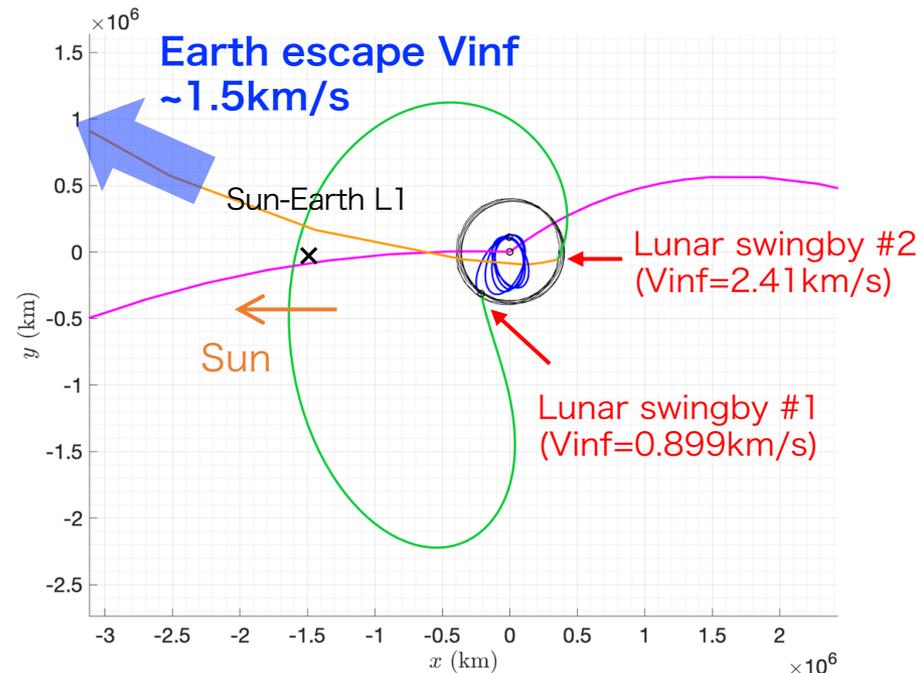
Launch & Separation



Exploiting solar gravity (tidal forces) can **increase the V_{inf} with respect to the moon!!**

In this case, the maximum V_{inf} w.r.t. the Earth is about 1.5km/s.

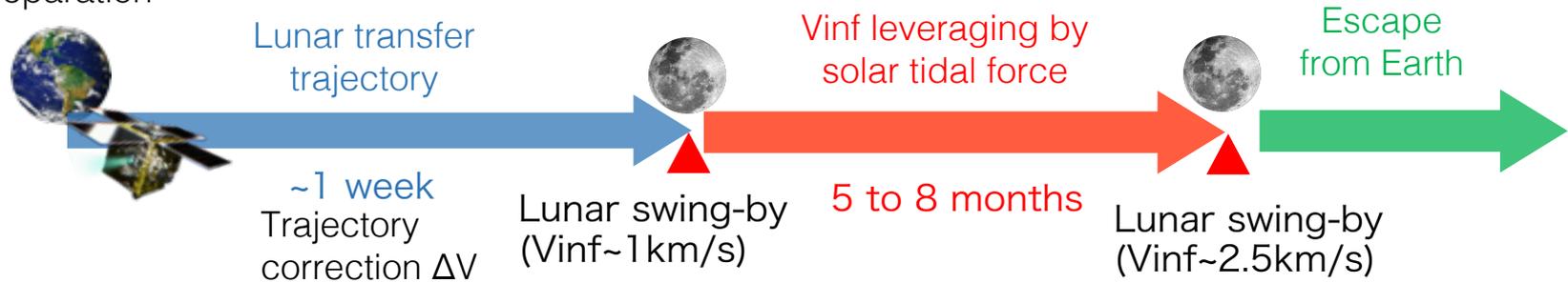
The interplanetary trajectory can be design under two-body problem (patched conics) with $V_{inf} < 1.5\text{km/s}$.



DESTINY+ Trajectory
(Ozaki, et al., 2019)

4) To Interplanetary Space (Mars, etc)

Launch & Separation



Reachable Planet for each Earth departure V_{inf}

(under Hohmann transfer assumption)

Planet	Earth departure V_{inf} , km/s
Mercury	7.53
Venus	2.50
Asteroid	Depending on the body
Mars	2.94
Jupiter	8.49
Saturn	10.29

We cannot reach most of them with $V_{inf}=1.5\text{km/s}$

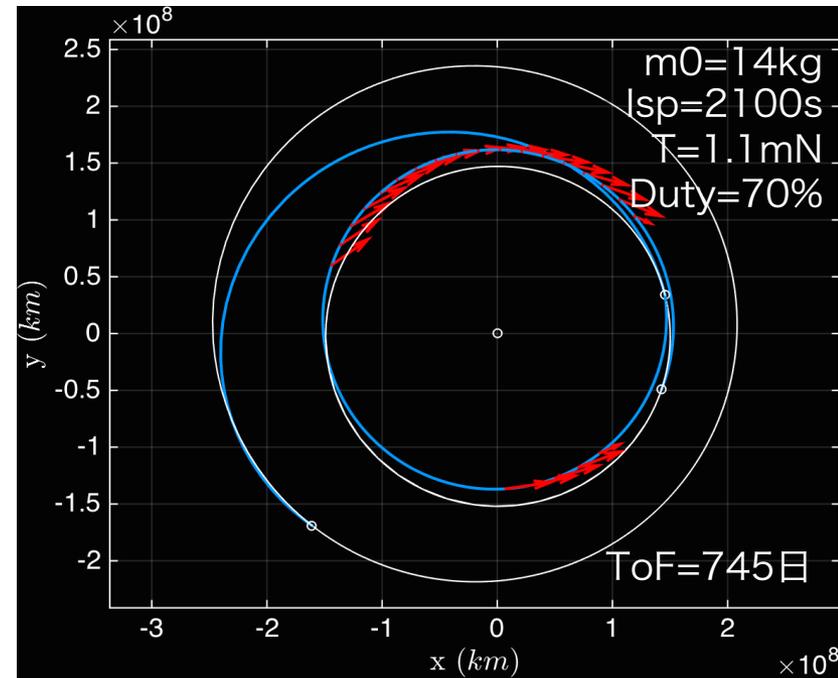
4) To Interplanetary Space (Mars, etc)

Launch & Separation



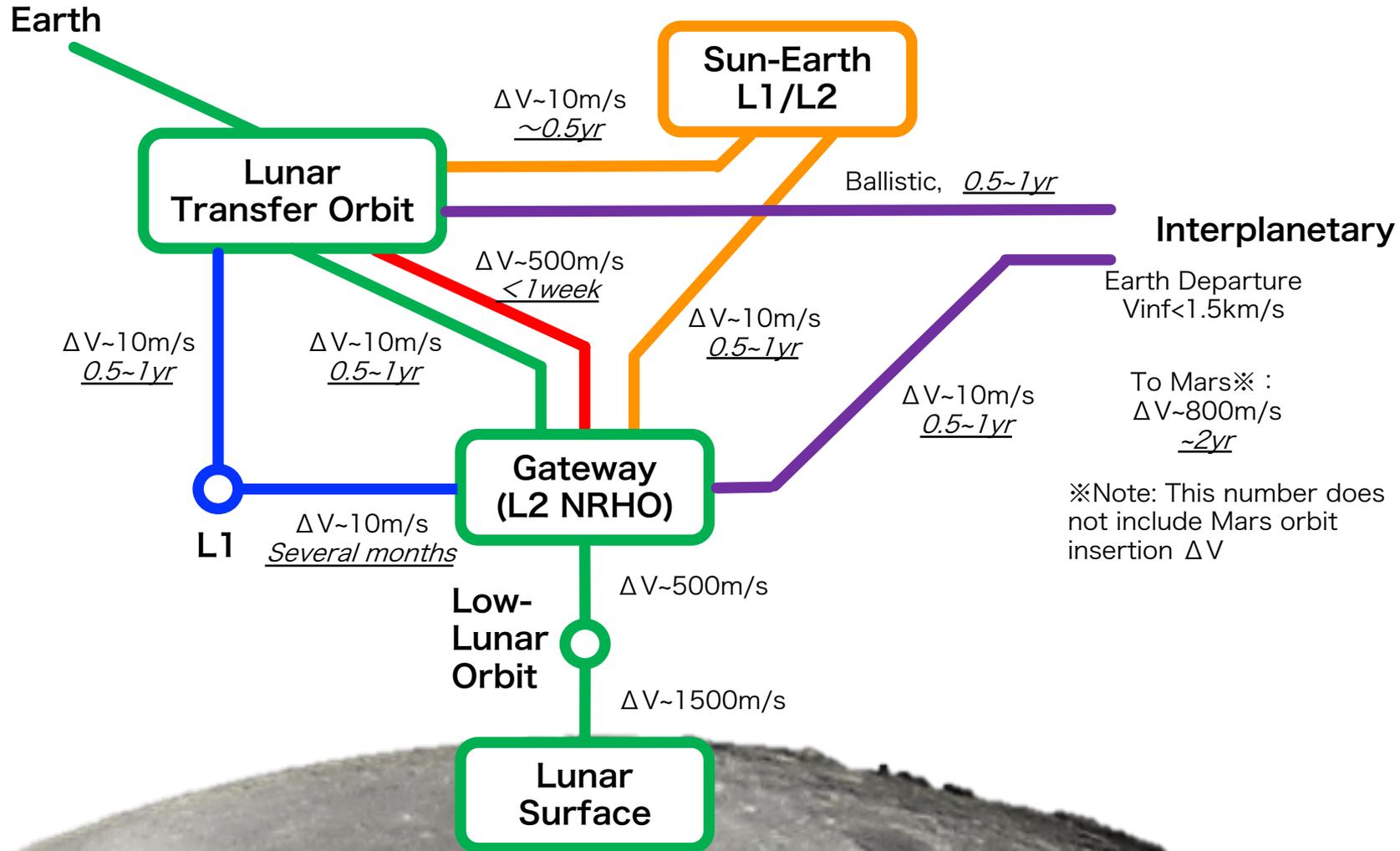
V-infinity leveraging transfer (or $\Delta V_{\text{VEGA/EDVEGA}}$) can effectively increase the Earth departure V_{inf} .

If we want to increase V_{inf} from 1.5km/s to 3.0km/s (reachable to Mars), **we need 0.83km/s ΔV (about half of V_{inf} increment)**



V-Infinity Leveraging Transfer to Mars

Gateway Metro Map



Possible Small Sat Mission Utilizing LOP-G

When we assume that we can deliver 6U CubeSat to LOP-G (or SLS/Artemis-1 like trajectory), the following missions are possible.

Target	Possible Using Current Technology	Challenging, Could Be Possible in the Future
Moon	<u>Moon orbiter</u>	<u>(Soft?) Landing</u>
Asteroid	<u>Flyby to NEAs</u>	<u>Rendezvous to NEAs,</u> Exploration to main belt asteroids
Lagrange Points	<u>Earth-Moon halo,</u> <u>Sun-Earth halo</u>	—
Mars, Venus	—	<u>Flyby exploration,</u> Orbiter? Lander?
Outer Planet	—	Dependent exploration (Stand alone mission could be possible if innovative technologies are developed)

Bold: Possible missions by SLS, Artemis-2

Summary

- What is the Lunar Gateway?
- Which orbit can the spacecraft transfer from lunar transfer orbit (Gateway construction opportunities)
- Which type of mission can the small spacecraft do by utilizing the Gateway opportunity.